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	L) IJP(c) JD/AT CODE: UR/0181/66/008/002/0424/0427
AUTHOR: Vakulenko, O. V.; Lisitsa, M. P.	5.5 54
RG: Kiev State University im. T. G. Shevchen versitet)	355
ITLE: Absorption of thermally excited carries	sin germanium
OURCE: Fizika tverdogo tela, v. 8, no. 2, 196	
OPIC TAGS: semiconductor carrier, thermal exc bsorption spectrum	citation, germanium semiconductor,
BSTRACT: The authors studied the absorption s 03, 490 and 530°K and that of heavily doped p- avelength range. The doped specimen was a gal $.2 \cdot 10^{-3}$ cm thick. Relatively pure natural spe as no absorption by free carriers at room temp and experimental data shows satisfactory agreem the $K_{13}$ band. The results indicate that the pr	Lium-activated germanium plate cimens were selected in which there erature. A comparison of theoretical
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of str both p of ind	uctural jab ure/and do lrect tran	on of carrier sorption by h ped specimens sitions and c gures, 1 table	oles in german . This type onsequently on	nium agrees w of excitation	ith the is prac	experimental tically inde	data foi pend <mark>ent</mark>
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<u>29919-66</u> EWT(1)/EWT(m)/T/EW ACC NR: AP6018038	SOURCE CODE: UR/0185/66/011/006/0653/0657
CC HAR AFOUL0030	
UTHOR: Vakulenko, O. V.; Lysy	tsya, M. PLisitsa, <u>M. P.; Zayets'</u> , V. DZayets, B
RG: Kiev State University im.	T. G. Shevchenko (Kyyivs'kyy derzhuniversytet); Institute
f Semiconductors, AN URSR, Kiyev	(Instytut napivprovidnykiv AN URSR)
	when implicited allicon
ITLE: Infrared spectrum of ne	eutron-irradiated silicon 19
OURCE: Ukrayins'kyy fizychnyy	zhurnal, v. 11, no. 6, 1966, 653-657
OPIC TAGS: neutron irradiatio	on, irradiation damage, irradiation effect, absorption
pectrum	
-	21
BSTRACT: An investigation was	made of the Si absorption spectrum in the range of or and after irradiation with neutrons. A polished
vavelengths from 1 to 25 µ beic	with a specific resistance of $\rho = 5$ ohm cm was
rradiated with neutrons with a	an average energy of 2 Mev and an integral flux in-
ensity of $10^{18}$ neutrons/cm <sup>2</sup> .	The temperature of the specimen in the reactor did not
xceed 100C. The investigation	a showed that the absorption band at $\lambda = 1.8 \mu$ is
pparently not linked to residu	al impurities whose concentration does not exceed of centers responsible for this absorption, calculated
sing the Kravets integral, sho	ould be approximately 2 x 1017 cm <sup>-3</sup> . Numerous defects-
i vacancies caused by irradiat	tion-play an important role in the creation of these
enters. It is possible also	that hydrogen or carbon, whose concentration in the
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he band result Cc-0.40 Si-vacar	at ) of 1 ev. ncies	the tra The a , creat	u. It nsition bsorpti ed by 1	is assur f of an e lon at λ = neutron i	definite ro med, howeve lectron fro = $12 \mu$ is rradiation ility. The toms. Ori	er, that to om the val explained , even at	ence band by the ex room tempe	to the istence rature reated	level of of Si-A- cannot pu when the	-centers roduce Si-
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, 41600-66 EWT(1)/EWT(m)/T/EWF(t)/ETI INP(c) ID	
ACC NR: AF6018528 SOURCE CODE: UR/0181/66/008/006/1698/1701	
AUTHOR: Vakulenko, O. V.; Lisitsa, M. P.; Kononets, Ya. F.	
ORG: Kiev State University im. T. G. Shevchenko (Kiyevskiy gosudarstvennyy univer-	
sitet)	
TITLE: Infrared absorption by carriers in lead sulfide	
SOURCE: Fizika tverdogo tela, v. 8, no. 6, 1966, 1698-1701	
TOPIC TAGS: lead compound, sulfide, ir absorption, electron density, Hall effect, absorption edge, carrier scattering	
ABSTRACT: The absorption spectrum of PbS was investigated in the range $\lambda = 3 - 15 \mu$ at temperatures 293 and 100K. The measurements were made on a small single crystal (0.22 mm thickness). The electron density necessary for comparison with theory was obtained from Hall-effect measurement and was found to be $2 \times 10^{17}$ cm <sup>-3</sup> at room tem- perature. The spectra exhibit a characteristic shift of the absorption edge towards longer wavelengths with decreasing temperature, and also a decrease in the absorption by the free carriers. The values of the absorption coefficient at the minimum of the absorption curve ( $\approx 18$ cm <sup>-1</sup> ) was found to be independent of the temperature. After illumination of this background, which is apparently connected with mechanical de- fects, the coefficient of absorption by the free carriers is found to be proportional to $\lambda^{2.8\pm0.2}$ , accurate to within 20%. Arguments are presented to show that the ab- sorption by the free carriers in PbS is not due to the impurity scattering mechanism,	
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ZYUZIN, Ivan Ivanovich; VAKULENKO, Sergey Mikhaylovich; SARANTSEV, Yu.S., red.

> [Organization and technology of the repair of freight cars; work practices of the Taiga Station depot of the Western Siberia Railroad] Organizatsiia i tekhnologiia remonta gruzovykh vagonov; oryt raboty vagonnogo depo st. Taiga Zapadno-Sibirskoi dorogi. Moskva, Transport, 1964. 74 p. (MIRA 17:9)

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·米亚市地区和西洋和市场和市场和市场。 VAKULENKO, S.M. Overall mechanization as the base of advanced technological processes. Zhel. dor. transp. 46 no.4:78-22 Ap 164. (MIRA 17:6) 1. Nachal'nik vagonnogo depo stantsii Tayga Zapadno-Sihirskoy dorogi. 

KOVALEV, M.M.; VAKULENKO, S.N.

In a final sector of the secto

State of the blood coagulation system in various force of goiter. Probl. endok. i gorm. 11 no.2:22-27 Mr-Ap '65. (MTRA 18:7)

1. Kafedra gospital'noy khirurgil (zav. - prof. M.M.Koralev) Kiyevskogo meditsinkogo instituta i 3-ye khirurgicheakeye otdeleniye klinicheskoy bol'nitsy im. Oktyabrskoy revolutsii (glavnyy vrach V.F.Alekseyev).

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VARU:NNKO, V. Flat collar on hydraulic tools for the removal of propallers. Mor. (MIRE 11:4) flot 18 no.3:23 Mr '58. 1. Vedushchiy inzhener TSentral'nogo konstruktorskogo byuro sudostroitel'noy promyshlennosti. (Shipa--Maintenance and repair)

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ZHUKOV, A.I.; RAFADTORV, YG.I.; TARULINEO, V.A.

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Separation of thorium and uranium (VI) on NE-1 resin. Lar. prikl. khim. 38 no.1:43-47 Ja '65.

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VAKULENKO, V.D.

Improving the technology for cutting out ballast sections in track overhauling. Put' 1 put. khoz. 8 no.5:20-21 My '64. (MIRA 17:6) l. Glavnyy inzh. putevoy mashinnoy stantsii No.121, stantsiya Novograd-Volynskiy 1, Yugo-Zapadnoy dorogi.

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KOVALMNKO, Yevgeniy Ivanovich; VAKULMNKO, V.P., redaktor; MAKAROVA, A.N., tekhnicheskiy redaktor. [Labor organization and discipline on collective farms] Organizatelia i distsiplina truda v kelkhese. Meskva, Ges.izd-ve iurid. lit-ry. 1955. 62 p. (Cellective farms) (MLRA 9:5) 

in the second STATISTICS AND A STATISTICS IN A At the destances KAZANTSEV, Nikolay Dmitriyevich; VAKULENKO, V.P., red.; SHCHEDRIMA, N.L., tekhn.red. [The charter of the agricultural artel] Ob ustave sel'skokhosiaistvennoi arteli. Izd.2-oe, ispr.i dop. Moskva, Gos.izd-vo iurid. (MIRA 10:12) lit-ry, 1957. 65 p. (Collective farms) 

PANKRATOV, Ivan Ferisanovich; VAKULENKO, V.P., red.; KOSAREVA, Ye.N., tekhn.red. [Legal forms of the responsibility of collective-farm officials] Pravovye fermy etvetstvennesti delzhnestnykh lits kolkhozov. Moskva, Ges.izd-ve iurid.lit-ry, 1959. 198 p. (MIRA 13:6) (Cellective farms -- Officials and employees) 



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[Biology and cultivation of tulips] Biologiia i kul'tura tiul'panov. [Moskva] 1959. 5 p. (Akademiia kommunal'nogo khoziaistva, Informatsionnoe pis'mo, no.5). (MIRA 16:8)

1. Sektor ozeleneniya Akademii kommunal'nogo khozyaystva (for Timofeyeva).

(Tulips)

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THE WAR LATE PROPERTY AND REAR REMEMBER AND A VERY IN WATTY OF

VAKULENKO V.V.

State of the production of flower seeds and measures for its improvement in the R.S.F.S.R. Trudy Bot.inst.Ser.6 no.7:415-418 (MIRA 13:4) 159.

1. Akademiya kommunal'nogo khosyayatva RSFSR im. K.D.Pamfilova, Moskva.

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[Annual flowers] Odnoletnie tsvetochnye rasteniia. Sost.V.V. Vakulenko i T.M.Aleinikova. Moskva, Izd-vo M-va sel'khoz.RSFSR, 1961. 259 p. (MIRA 14:4)

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VINOGRADOV, K.A.; ZEMLYANITSKIY, L.T.; NOVOZHILOVA, V.A.[deceased]; LUNEVA, Z.S.; <u>VAKULENKO, V.V</u>,; GALAKTIONOV, I.I.; ALEKSEYENKO, L.V.; NERONOVA, M.D., red.; KHENOKH, F.M., tekhn. red.

> [Care of urban plantings] Ukhod za gorodskimi nasazhdeniiami. Moskva, Izd-vo Kommun. khos.RSFSR, 1963. 89 p. (MIRA 16:7)

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Study of the solvent flooding process. Izv. AN Azerb. SSR. Ser. geol.-geog. nauk nc.2:104-111 '65. (MIRA 18 (MIRA 18:8)

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Investigating the strain of the support frame of the 5-4 selfpropelled combine by static tests with the electric strain guage. Nauk, pratsi Inst. lyv. vyrob. AN URSR 7:54-80 '59. (MIRA 14:1) (Combines (Agricultural machinery)) (Strains and stresses)

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Investigating the dynamic strain of the support frame elements of the S-4 self-propelled combine by electric strain gauging. Eauk. pratsi Inst. lyv. vyrob. AN URSE 7:81-107 '59. (MIRA 14:1) (Combines (Agricultural machinery)) (Strains and stresses)

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So: Knizhnaya Letopis', No. 18, 1956

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"Forest Growth Conditions on the Sands in the Lower Course of the Medvedista and Khopr Rivers and the Pinc Crops on Them." Cand Agr Sci, Saratov Agricultural Inst, Min Grops USSR, Saratov, 1954. (KL, No 7, Feb 55)

SO: Sum. No. 631, 25 Aug 55- Survey of Scientific and Technical Dissertations Defended at USSR Higher Educational Institutions (14)

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THE REPORT OF THE

VAKULIN, A.A.; V'YUNOV, S.F.; GORIN, T.I.; IVASHCHENKO, P.S.; KOMOVA, A.G.; KORNEYEV, V.A.; KOROSTELEVA, M.Ye.; LOBACHEV, A.Ye.; LASHMANOV, I.Ya.; MALYCHENKO, V.V.; MOROZOVA, A.M.; PANSHIN, I.A.; PROSVIROV, A.S.; ROZHKOVA, M.V.; YUROVA, N.F.; FEDORENKO, V.P.; TSEKHMISTRENKO, P.Ye.; SHEVCHENKO, I.S.; FEDOROV, N.A., red.; IZHBOLDINA, S.I., tekhn.red.

[Brief menual on the cultivation of fruits, berries, and grapes and the management of nurseries in Stalingrad Province] Kratkii spravochnik po plodovo-iagodnym kulturam, vinogradu i pitomnikam dlia Stalingradskoi oblasti. Stalingrad, Stalingradskoe knizhnoe izd-vo, 1960. 215 p. (MIRA 14:3)

1. Stalingred (Province) Upravleniye sel'skogo khozyaystva. (Stalingrad Province--Fruit culture)

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Main sand-fixing plants in the deserts and semidesets of Inner Mongolia (Chinese People's Republic). Bot.zhur. 47 no.11:1680-1684 N '62. (MIRA 16:1)

1. Volgogradskiy sel'skokhozyaystvennyy institut. (Mongolia-Desert flora)

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**全部的部位**法

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PLYUSHCH, A.M.; VAKULIN, A.N. dente statues trans. Experience in the use of solvents to increase petroleum recovery. Nefteprom. delo no.4:36-39 163. (MIRA 17:8) 

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AMMENKOV, V.A., kend. t-kim. rauk: YAKULIM, Y.S., Inzh.; HIGHENKO, 1.G.; ROCOZHÍN, P.A. Threaded packing of a molding vulcanizer. Khim. i neft. massimostr. no.2:40 F 165. (MIRA 18:4) 





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"APPROVED FOR RELEASE: 08/31/2001 CIA-RDP86-00513R001858420003-0 VAKULOV, Nikolay-Fedorovich; GAYDENKO, V.M., retsenzent; KOSUL'NIKOV, N.K., retsenzent; MAKRUSHINA, A.N., red.izd-va; RIDNAYA, I.V., tekhn. red. [Diesel and electric crane operator's manual]Posobie kranovshchiku dizel'nogo i elektricheskogo krana. Moskva, <sup>I</sup>zd-vo "Rechnoi transport," 1961. 202 p. (MIRA 15:12) "Rechnoi transport," 1961. 202 p. (Cranes, derricks, etc.) \$ in the second second

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•	sov/20-125-2-16/64
24( 7) ,24( 8) AUTHOR <b>S</b> :	Vernov, S.N., Corresponding Member, AS USSE, Chudakov, A.Ye., Vakulov, P.V., Logachev, Yu.1.
TITLE:	Investigation of Terrestrial Corpuscular Radiation and of Cosmio Rays During the Flight of a Cosmio Rocket (Izucheniye zemnogo korpuskulyarnogo izlucheniya i kosmicheskikh luchey pri relete kosmicheskoy rakety)
PERIODICAL:	Doklady Akademii nauk SSSR, 1959, Vol 125, Nr 2, pp 304 - 307 (USSR)
ABSTRACT:	The rocket launched on January 2, 1959 in the direction of the moon had apparatus for recording cosmic- and terrestrial corpuscular radiation on board. By the latter the authors mean the fluxes of fast charged particles in great altitudes, for which the terrestrial magnetic field is a $\infty$ -salled "magnetic trap". The particles were recorded by 2 Geiger-counters and 2 sointillation-counters. The first apparatus, with scintillation counter, was a constructive further-development of the device which the authors had built into the third Soviet Sputnik. A cylindrical sodium-iodide crystal served as a detector. The authors, above all, described the results obtained by the preliminary evaluation of the data ascertained in altitudes of from 8000 to 150000 km (from the center of the earth). A schematical drawing shows the trajectory of the rocket with
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Investigation of Radiation and of Terrestrial Corpuscilar SOV/20-125-2-16/6L Cosmic Rays During the Flight of a Cosmic Rocket

respect to the terrestrial magnetic field. The intensity maximum is  $\sim 26000$  km from the center of the earth. At a distance of 55000 km the intensity of terrestrial corpuscular radiation becomes practically equal to zero, and the remaining ionization in this distance is entirely due to cosmic radiation. According to the authors' opinion the particles oscillate along the lines of force symmetrically to the equatorial plane. The increase of intensity along a given line of force in the transition from low to high altitudes serves as an experimental proof for this assumption. The particle flux is directed not only towards one side, and, in any case, the predominant part of the particles undergoes complete reflection when approaching the earth, and is therefore subjected to oscillations from one hemisphere to the other. The trajectory of the rocket nowhere . Actually, the apparatus intersects the so-called internal zone. built into the cosmic rocket in no range of their trajectory record particles of high energy which are characteristic of the inner zone. On the other hand, the composition of radiation is very similar to that observed by means of the third Sputnik in polar regions. Next, the composition of radiation in the outer zone with

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Investigation of Radiation and of

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Terrestrial Corpuscular SOV/20-125-2-16/64 Cosmic Rays During the Flight of a Cosmic Rocket

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high intensity is dealt with. In the center of the outer zone, where particle density is the greatest, the effective energy of electrons is minimal. In conclusion, cosmic radiation is dealt with. Beginning with a distance of 66000 km, the intensity of all components remains constant. The strict constants of all components at distances of from 66000 to 150000 km indicates the existence of a radiation upon which the terrestrial magnetic field exercises no influence. Therefore, either the terrestrial magnetic field vanishes at a distance of 10 earth-radii, or there are no particles with momenta of

 $1.5.10^8$  to  $4.10^7$  ev/c in interplanetary space. The energy-flux of the photons is very low and contributes partly nothing to ionization. There are 2 figures and 4 Soviet references.

SUBMITTED: February 25, 1959

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		AVAILABLE: LADRANY OF Congress Chird 3/3	Thirrarkly, J. M., and G. V. <u>Ourton</u> . Change of the Albeds of the <u>Fort Artificial</u> Kirth Shellits Busulting From the Action of Th Sidemal Photons		Boolars <u>St. I.</u> Construing the Problem of the Portation of 50° in 60	and the map of minimum and an an open frobless of Control in Baubanhah, B. Y., and B. B. Thint'. Some Probless of Control in Divergions and Speci	Ernoson, L. V., V. I. Lercher, L. A. Maoresor, and M. It. Jenkin. Investigation of Could Builailed in the Fiight of the Deroil Space- Boche to the Moon Bacarma, T. S. Realls of the Investigation of Meteoric Buistance Unitable First Fouriers Annuted in Space Reputs	Vernor, S. T., A. Te, Chudaker, <u>P. V. Vabulor, Du. T. Logechav</u> , and T. U. Himilayev, Wallation Measurement in the Flight of the Second Space Rocket	portrained in the solution of 10 articles deals with ynobless of scilling methods, magnetic security relations to relation of energy to security the solution of the solutions. Bo presentities are mailtined. Betweener accounty sees of the articles. Bo presentities are mailtined. Betweener accounty of a the articles. Betweener accounty of the articles. Betweener accounts of the articles. Because the second (Seriet) Space to be articles. Because the second (Seriet) Space to be articles.	O. M. Gus'horm. NYNCOE: The boolist is intended for scientists and engineering and scientific personnel working in the field of space trevel and satellite flight.	Likussiwammyre spuidiki senli, vyp. 5 (Artificial Earth Exhilites, No. 5) Rescor, Lid-vo AN SKER, 73 p. Ernska ellp inserted. 7,000 copies printed. Resp. M.: L. V. Eurosovej Ed. of Publishing Bouset M. L. Predicing Toch. Ed.:	FIRST I NOOT EXTRATION SOV/1222		
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3,2000	<b>\$/020/60/130/03/009/06</b> 5
-29 (2), -29 (5 AUTHORS:	Vernov, S. N., Corresponding B014/B014
	Nember of the AS booky . Yu. I., Nikolayev, A. G.
TITLE:	Radiation Measurement During the Flight of the Second Cossid
- PERIODICAL:	Rocket v Doklady Akademii nauk SSSR, 1960, Vol 130, Nr 3, pp 517 - 520
	(USSR) The equipment of the interplanetary rocket launched on The equipment of the interplanetary rocket launched on
ABSTRACT:	
	are described in detailed ion of data obtained for the lange
	from 9,000 to 120,000 an and Figure 1 illustrates the tra-
Card $1/3$	jectories of the first and second interplaneticly measure- ferred to the terrestrial magnetic field. Ionization measure-
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	enter and a second second second second and second and second second second second second second second second

Rediation Measurement During the Flight of the S/020/60/130/03/009/065 Becond Cosmic Rocket Bolt4/B014 ments are also graphically represented in figure 1. It is noted that the shift of the ionization maximum between the two measurements was not caused by the slight difference of the trajectories of the two rockets. The radiation belt is most probably deformed by streams of solar corpuscles. This assumption seems to be confirmed by a comparison with the re- sults of measurements performed by the American rocket Pioneer III. The energy-flux density of electrons of more than 5 Mev or of protons of more than 30 Mev is said to be 1 particle/cm <sup>2</sup> .sec. Furthermore, a radiation was detected which consisted of electrons having an energy of the order of 10 <sup>6</sup> ev, or of protons of an energy of about 10 Mev. The first possibili- ty is considered to be more probable. This electron flux is said to be 5.10 <sup>5</sup> particles/cm <sup>2</sup> .sec. The existence of electron fluxes having an energy of between 20 and 50 kev (flux: 10 <sup>10</sup> particles/cm <sup>2</sup> .sec), which had already been detected by		67908
noted that the shift of the ionization maximum between the two measurements was not caused by the slight difference of the trajectories of the two rockets. The radiation belt is most probably deformed by streams of solar corpuscles. This assumption seems to be confirmed by a comparison with the re- sults of measurements performed by the American rocket Pioneer III. The energy-flux density of electrons of more than 5 Mev or of protons of more than 30 Mev is said to be 1 particle/cm <sup>2</sup> .sec. Furthermore, a radiation was detected which consisted of electrons having an energy of the order of 10 <sup>6</sup> ev, or of protons of an energy of about 10 Mev. The first possibili- ty is considered to be more probable. This electron flux is said to be 5.10 <sup>5</sup> particles/cm <sup>2</sup> .sec. The existence of electron fluxes having an energy of between 20 and 50 kev (flux:		s/020/60/130/03/009/065 B014/B014
said to be 5.10 <sup>5</sup> particles/cm <sup>2</sup> .sec. The existence of electron fluxes having an energy of between 20 and 50 kev (flux:	noted that the shift of the ioniz two measurements was not caused b the trajectories of the two rocke most probably deformed by streams assumption seems to be confirmed sults of measurements performed b Pioneer III. The energy-flux dens 5 Mev or of protons of more than 1 particle/cm <sup>2</sup> .sec. Furthermore, consisted of electrons having an or of protons of an energy of abo	ation maximum between the y the slight difference of ts. The radiation belt is of solar corpuscles. This by a comparison with the re- y the American rocket ity of electrons of more than 30 Mev is said to be a radiation was detected which energy of the order of 10 <sup>6</sup> ev, ut 10 Mev. The first possibili-
10 particles/cm <sup>2</sup> .sec), which had already been detected by	said to be 5.10 <sup>5</sup> particles/cm <sup>2</sup> .sec fluxes having an energy of betwee	The existence of electron n 20 and 50 kev (flux:
the first intercontinental rocket, were proven again. Thus, Card 2/3 two essential groups of particle fluxes were found: electrons	the first intercontinental rocket	were proven again. Thus,

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67908 Radiation Measurement During the Flight of the s/020/60/130/03/009/065 Second Cosmic Rocket B014/B014 of about 20 kev and 10<sup>6</sup> ev electrons. The energy of the first group is close to the mean energy of the solar corpuscular radiation and allows to assume the existence of a thermodynamic equilibrium between protons and electrons on their penetration into the terrestrial magnetic field. It is pointed out that the electron momenta of the second group are close to the proton momenta of corpuscular radiation and to the momenta of the electrons arising from the decay of the reflected neutrons. The existence of a lunar radiation belt could not be proven. Constant radiation intensity was measured at a distance of 70,000 km from the Earth. There are 2 figures, 1 table, and 2 references, 1 of which is Soviet. SUBMITTED: November 20, 1959 Card 3/3

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· ... • 34350 s/203/61/001/006/002/021 9,6150 D055/D113 3.2420 (1049, 1482) Vernov, S.N.; Chudakov, A.Ye.; Vakulov, P.V.; Gorchakov, Ye.V.; AUTHORS: Logachev, Yu.I. Radiation measurements in the outer radiation belt on February 12, 1961, during the rocket flight towards Venus TITLE: PERIODICAL: Geomagnetizm i aeronomiya, vol 1, no 6, 1961, 872-874 TEXT: The article deals with data on the Earth's outer radiation belt collected when the Earth-Venus rocket launched on February 12, 1961, was 30,000 - 45,000 km from the Earth's center. The special equipment installed in the hermetic container consisted of a scintillation counter and an CTC-5(STS-5) gas-discharge counter. The distribution of matter around the NaJ(T1) crystal and the gas-discharge counter is shown in a table. By reducing the dimensions of the crystal and increasing the resolving power of the electronic system of the counter, the radiation intensity in the belt was correctly registered. Fig. 1 shows the overload characteristics for the counting channels of the scintillation (1) and gas-discharge (2) counters. Card 1/6

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Radiation measurements ...

These channels could register up to  $10^6$  and  $10^5$  pulsations/sec. respectively. To penetrate the crystal of the scintillation counter and the working volume of the gas-discharge counter, electrons must have an energy of > 3 Mev, protons - an energy of > 32 Mev and the bremsstrahlung quanta - an energy of > 30 kev. Curves on fig. 2 represent the counting speed of the scintillation counter (1), that of the gas-discharge counter after corrections were made according to the curves in fig. 1 (3) and the energy release in the crystal in relation to the distance from the Earth's center (2). As all three curves were more or less parallel, the mean energy release in the crystal for one reading of the scintillation counter was 130 kev and remained constant between 32,000 and 40,000 km and the mean energy of the bremsstrahlung quanta did not vary with distance. The constancy of the meanenergy release showed that no great changes occurred in the spectrum of electrons of the outer radiation belt. A diagram (fig. 3) shows the paths of the interplanetary rocket (curve 1) and those of another three Soviet rockets (curve 2). A comparison of radiation and ionization data concerning the interplanetary rocket and the space rockets no. 1 and 2, showed that the outer radiation belt was stable for a period of 2 years when no magnetic perturbances were recorded. However, this period was not long enough to

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Radiation measurements ...

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evaluate solar effects on the outer belt, but could be taken as an indication of the absence of such an effect. The space rocket no 3 was launched during a moderate magnetic storm (the change in the vertical and horizontal components of the terrestrial magnetic field was about 250 and 150 y respectively) The external side of the belt was not measured, but the total energy release in the crystal during the entire flight coincided with that calculated for the rocket no 1 and was 1.5 times less than that of the rocket no 2, i.e. no changes occurred in the mean state of the outer zone during the flight of the rocket no 3 during a moderate magnetic storm. Since measurements were started a few hours after the beginning of a magnetic storm, the radiation intensity in the belt had not yet decreased. On the other hand, it is also possible that not all magnetic storms cause the radiation intensity of the Earth's outer radiation belt to decrease. There are 4 figures, 1 table and 3 non-Soviet references. The three Englishlanguage references are: W.H. Hess, J. Geophys. Res., 1960, 65, no 10, 3107; P. Rothwell, C.E. McIlwain. J. Geophys. Res., 1960, 65, no. 3, 799; R.L. Arnoldy, R.A. Hoffman, J.R. Winckler, J. Geophys.Res., 1960, 65, no 5, 1361.

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34352 s/203/61/001/006/004/021 D055/D113 9.6150 Vakulov, P.V.; Goryunov, N.N.; Logachev, Yu.I., and AUTHORS: Sosnovets, E.N. Radiation registered during the flights of Soviet artificial TITLE: satellites and space rockets Geomagnetizm i aeronomiya, v.l , no.6, 1961, 880-887 PERIODICAL: TEXT: Methods of registering radiation, based on the use of scintillation and gas-discharge counters and applied in Soviet artificial satellites and space rockets, are described. The registration of weak currents (up to  $10^{-10}$ ) with the aid of a relaxation oscillator on a neon tube is described. The scintillation counters measured the number of particles releasing more energy in the crystal than that determined by the thresholds of the counting devices. Ionization caused by radiation in the entire crystal was also measured. The gas-discharge counters registered charged particles and  $\chi$  radiation to an accuracy of  $\sim 1\%$ . The counters were located behind screens of various materials to facilitate the analysis of radiation according to Card 1/3

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## CIA-RDP86-00513R001858420003-0

Radiation registered ...

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penetration. The devices with the counters were located both inside and outside the container with scientific apparatus. For economy the photomultipliers in the counters were fed without a divider by leads from a high voltage battery direct to the electrodes. Ionization was determined from the currents of the anode and seventh dynode. By using two channels, these currents could be compared in order to estimate how much of the energy produced in the crystal resulted from saturation of the anode current during intense scintillation in the crystal. By this means comparatively highenergy particles could be detected in the inner zone during tests with the third artificial Earth satellite. The use of a single scintillation counter to measure many parameters permitted the weight and size of the device to be reduced but required careful selection of photomultipliers, which had to satisfy the following requirements: (1) there must not be more than one sound impulse per 10 sec. corresponding to energy liberation of above 30 kev in an NaJ(T1) crystal and there must be practically no impulses corresponding to energy liberation of > 300 kev; (2) leakage current of the seventh dynode  $\leq 1 \cdot 10^{-10}$  a; (3) anode dark current  $\leq 1 \cdot 10^{-8}$  a; (4) leakage current of the other dynodes of the intervals  $\leq 1 \cdot 10^{-7}$  a. The CTC-5(STS-5)

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Radiation registered ...

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gas-discharge counters used were small and had a low operating voltage  $(\sim 400 \text{ v})$  and a thin wall, which facilitated the recording of low-energy particles. The electronic circuits operating on semi-conductor elements and the calibration method are described in detail. There are 9 figures and 5 Soviet references.

ASSOCIATION: Moskovskiy gosudarstvennyy universitet im. M.V. Lomonosova. Institut yadernoy fiziki (Moscow State University imeni M.V. Lomonosov. Institute of Nuclear Physics).

SUBMITTED: October 12, 1961

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or housed with the second states and the second states and the second states and the second states and second s 1. 89610 s/020/60/136/002/013/034 B019/B056 9,9130 (1041, 1046, 1060) Vernov, S. N., Corresponding Member of the AS USSR, AUTHORS: Chudakov, A. Ye., Vakulov, P. V., Gorchakov, Ye. V., Logachev, Yu. I., and Nikolayev, A. G. Radiation Measurements During the Flight of the Third Cosmic TITLE: Rocket 1961 Doklady Akademii nauk SSSR, 1960, Vol. 136, No. 2, pp. 322-324 PERIODICAL: TEXT: The third cosmic rocket launched on October 4, 1959 contained a scintillation counter and three gas discharge counters. All gas discharge counters had a wall strength of 50 mg/cm<sup>2</sup> steel sheets and were, in addition, surrounded by several shields. Counter I had a shield made from 3 mm lead + 1 mm aluminum with a counter window of 0.28  $cm^2$ , which was closed by a 0.2 mm thick aluminum sheet. Counter II had the same shield, but without counter window, and counter III was in an aluminum container made from 2.5 mm thick aluminum. The scintillation counter recorded the ionization of the crystal (NaI) and the counting rate. Preliminary results of evaluation of the instrument readings are given from the time from Card 1/2

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Radiation Measurements During the Flight of the Third Cosmic Rocket s/020/60/136/002/013/034 B019/B056

October 4, 1959, to October 18, 1959. The trajectory of the rocket was in practical agreement with that of the first and second cosmic rocket. From a comparison of the readings of the various counters, the authors conclude that the intensities of the particles recorded by the instruments depend on the absorption in the container walls. Measurements in the interplanetary space showed that the cosmic radiation on the boundary of the terrestrial magnetic field is very strong; only individual small fluctuations were recorded. Finally, the agreement existing between the recorded intensities and those of a monitor are dealt with. From these considerations the authors draw the conclusion that the weak variations in the time from October 4 to October 18 are in connection with the variations of the magnetic fields in the solar system and the interactions among the latter are connected with cosmic radiations. There are 1 figure, 1 table, and 3 Soviet references.

SUBMITTED: October 26, 1960

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APPROVED FOR RELEASE: 08/31/2001

## CIA-RDP86-00513R001858420003-0

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AUTHORS:

<u>Vakulov, P. V.</u>, Vernov, S. N., Gorchakov, Ye. V., Logachev, Yu. I., Nesterov, V. Ye., Nikolayev, A. G., Pisarenko, N. F., Savenko, I. A., Chudakov, A. Ye., and Shavrin, P. I.

TITLE: Radiation studies during the flights of satellites, spaceships and rockets

PERIODICAL: Akademiya nauk SSSR. Izvestiya. Seriya fizicheskaya, v. 26, no. 6, 1962, 758-781

TEXT: This report deals with radiation measurements made by the second and the third Soviet spaceship, by the rocket launched toward the Venus on February 12, 1961, and by the third Soviet earth satellite (August 15, 1958). The spaceships were equipped with scintillation counters, gas discharge counters and elements for storing data through 24 hours. The northern and southern zones of increased radiation intensity are undoubtedly linked by the lines of force of the geomagnetic field. The increased radiation intensity is due to electrons of the outer radiation belt, slowed down in the jacket of the spaceship. The Card 1/3

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## CIA-RDP86-00513R001858420003-0

Radiation studies during the flights ...

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boundaries of this belt were determined more accurately by the lower orbiting Soviet spaceship. At 16 hours after the chromosphere flare of June 17, 1958 had vanished but still a few hours before the magnetic storm, charged particle intensity increased. The electron spectrum of the outer radiation belt does not change much at an altitude of 32,000-40,000 km, nor did the magnetic storm which occurred during the flight of the third Soviet spaceship have any substantial effect on the outer radiation belt. Except for a few percent, the proton intensity of the inner radiation belt remained constant during the three weeks' flight of the third Soviet satellite. The increased radiation intensity over the Brazilian anomaly, observed on board of the second spaceship at an altitude of 320 km, was due to the inner radiation belt. In this anomaly, the proton component of the inner radiation belt is predominant at small geomagnetic latitudes. The portion of X-rays increases with increasing latitude. A zone of lower bremsstrahlung intensity separates the outer from the inner radiation belt. This zone is practically absent in the region of the Brazilian anomaly. The equator of cosmic rays determined by the second and the third Soviet spaceship resembles remotely a sine curve running between 11° of northern and 11° of southern latitude.

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VAKULOV, P.V., VERNOV, S.N., GORCHAKOV, YE.V., LOGACHEV, YU.I. CHARAKHCHYAN, A.N., CHARAKHCHYAN, T.N., CHUDAKOV, A. YE.

Cosmic rays in the stratosphere and their correlation with solar activity.

Report to be submitted for Space Research Committee on COSPAR 6th plenary meeting Warsaw, Poland 11 June 63-



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JTHOR: <u>Vakulov, P. V.</u>	2
ITLE: Radiation study by the "Kosmos-17" satellite	FI I
OURCE: Vsesoyuznaya konferentsiya po fizike kosmicheskogo prostranstva. 965. Issledovaniya kosmicheskogo prostranstva (Space research); Trudy kom oscow, Izd-vo Nauka, 1965, 393-394	
OPIC TAGS: primary cosmic ray, satellite data analysis, radiation belt, harge counter, scintillation counter	
BSTRACT: The "Kosmos-17" carried equipment for studying emission from the adiation belts and primary cosmic radiation at altitudes up to 800 km. The sound and a gas discharge counter were included in this equipation counters and a gas mounted on the outer shell of the satel of the sound state of the sound state of the satel of the satel of the sound state of the sound sta	pment. One lite for
f the scintillation counters was mounted on the outer them to and 18 wo-threshold recording of electrons with energies greater than 50 and 18 rotons with energies greater than 600 kev. This same counter was used for any protons with energies greater than 5.4 and 8.5 Mev on two sep- hresholds. The second scintillation counter as well as the gas discharge	or record-
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	ACCESSION NR: AT5023609 were located under an alumin The data obtained by the sate recorded in radiation belts a and emission genetically asso the American high-altitude St belt. The average lifetimes shells. The absolute fluxes spectrum of trapped protons a zones are determined where th that these regions coincide to	and 2) emission outsid ociated with cosmic ra- tarfish explosion were of these electrons an of these electrons an and their intensities	ie of radiatic ays). Large of a recorded in re determined re found as W in the inner trapped radi	on belts (cosmic electron flows du the inner radiat for various magn ell as the energy belt. The spati ation. It is sho	rays e to ion etic
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L 1553-66 FSS-2/EWT(1)/FS(	(v)-3/FCC/EWA(d)/EWA(h)	TT/GS/GN	
ACCESSION NR: AT5023610		UR/0000/65/000	/000/0394/0405
AUTHOR: <u>Vernov, S. N.;</u> Chuć <u>Kuznetsov, S. N.;</u> Logachev, <u>Rubinshteyn</u> , <sup>47</sup> I. A.; <u>Stolpove</u>	Yu. I.; Nikolayev, A. skiy, V. G.; El'tekov,	1.: <u>Bosnovets, E.</u> V. A. <u>V.</u>	N. 55 84
TITLE: Geometric position a	and particle composition	of the earth's r	adiation belts
SOURCE: <u>Vsesoyuznaya konfer</u> 1965. Issledovaniya kosmich Moscow, Izd-vo Nauka, 1965,	heskogo prostranstva (Sj		
TOPIC TAGS: cosmic radiatio	on, <u>earth radiation belt</u> HMASS, 12-	t, cosmic ray, <u>Ele</u>	ktron 1, Elek-
ABSTRACT: An exhaustive stu satellites, which were launce Table 1 of the Enclosure. I passed their apogee at about radiation belt was thus cross return branch of the orbit.	ched on 30 January 196 The first orbits were po t 3 o'clock a.m. local tin ssed at about midnight a	4. Orbital data a ositioned so that de. The outer bou and again at about were shifted towa	re given in the satellites ndary of the 7-8-gim. on the rd the sunset:
Elektron-1, by o min, and El			
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## ACCESSION NR: AT5023610

tron-1 and -2 were equipped with similar instrumentation. In some cases, however, there were differences in energy thresholds. A chart summarizing all data shows the electron and proton fluxes of different energies in the equatorial plane and for comparison gives IMP-1 data. The following conclusions can be made from the chart: 1) A belt of artifically injected electrons exists at distances closest to the Earth's center The maximum of the belt in February 1964 was at L = 1.35. The flux of electrons with energy above 2 Mev at the maximum was about  $1 \times 10^7 \text{ cm}^{-2} \cdot \text{sec}^{-1} \cdot \text{ster}^{-1}$ . 2) The average directed flux of protons with an energy of 45-70 Mev at the maximum of the inner belt (L = 1.45) was about 1.5 x  $10^3$  cm<sup>-2</sup>·sec<sup>-1</sup>·ster<sup>-1</sup>. A change in the integral spectrum at proton energies above 50 Mev was observed at L = 2.2; the spectrum of these energies is in the process of hardening, which could be explained by the theory of albedo neutrons. 3) The spatial distribution of protons with an energy of one to several Mev differs from that of the electrons. There is a definite regularity in the distribution of protons according to their energies. The average directed flux of protons with an energy above 2 Mev was about 4.5 x 10<sup>5</sup> cm<sup>-2</sup>.sec<sup>-1</sup>. ster<sup>-1</sup> in the equatorial plane at L = 2.8. It appears that the majority of the protons in this energy range are created by transverse drift with respect to the magnetic field lines. 4) A belt of high-energy electrons was observed at L = 2.75. Its width at the equator was about 0.4 earth radii. The average directed flux of electrons above 6 Mey was about 10<sup>2</sup> cm<sup>-2</sup>, sec<sup>-1</sup>, ster<sup>-1</sup>. 5) A minimum of distribution Card

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	150 kev energy was obser		
	intensity shift is subject negligible magnitudes.		
positioned, on the av	erage, at $L = 4.8$ . The n	naximum altitude int	tensity shift in-
	-0.2 within a wide range		
	e at $L = 7 + 0.5$ . On the rved. The average direct		
of over 70 kev at the	maximum of the outer be.	It is about $5 \times 10^6$	cm <sup>-2</sup> .sec <sup>-1</sup> .ster <sup>-1</sup>
and can abance by you	a than an order of magnif	ude. The electron	energy spectrum
observed within the 7	0 to 600 key range is in	agreement with the	data of other re-
observed within the 7 searchers. The elect be softening, in comp		agreement with the	data of other re- reil Meviappears to Orig. art. has:
observed within the 7 searchers. The elect	0 to 600 key range is in rongenergy spectrum in th	agreement with the	data of other re-
observed within the 7 searchers. The elect be softening, in comp	0 to 600 key range is in rongenergy spectrum in th	agreement with the	data of other re- reil Meviappears to Orig. art. has:
observed within the 7 searchers. The elect be softening, in comp 11 figures: ASSOCIATION: none	0 to 600 key range is in ron energy spectrum in th arison with measurements	agreement with the	data of other re- vel'Mev appears to Orig. art. has: [FP
observed within the 7 searchers. The elect be softening, in comp 11 figures:	0 to 600 key range is in rongenergy spectrum in th	agreement with the	data of other re- reil Meviappears to Orig. art. has:
observed within the 7 searchers. The elect be softening, in comp 11 figures: ASSOCIATION: none	0 to 600 key range is in ron energy spectrum in th arison with measurements	agreement with the	data of other re- vel'Mev appears to Orig. art. has: [FP
observed within the 7 searchers. The elect be softening, in comp 11 figures: ASSOCIATION:	0 to 600 kev range is in ron energy spectrum in th arison with measurements ENCL: 01	agreement with the	data of other re- vel'Mev appears to Orig. art. has: [FP SUB CODE: AA, SV



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				$O_{i}$	
	Table	1. Orbital da	ta		
		Elektron-1 (low altitude)	Elektron-2 (high altitude)		
	Altitude, apogee	7, 140 km	68,200 km		
	Altitude, perigee	406 km	460 km		
	Orbital period	2 hr 48 min	22 hr 30 min		
	Inclination of orbital plane	61°	61°		ľ
	Period of rotation	40 sec	120 sec		
	1				
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ACCES	281-66    FSS-2/EWT(1)/FS(v)-3/FCC/EWA(d)/EWA(h)    T <sup>T</sup> /GS/GW      081-66    FSS-2/EWT(1)/FS(v)-3/FCC/EWA(d)/EWA(h)    T'/GS/GW      081-66    FSS-2/EWT(1)/FS(v)-3/FCC/EWA(d)/EWA(h)    T'/GS/GW      081-66    FSS-2/EWT(1)/FS(v)-3/FCC/EWA(d)/EWA(h)    F      081-66    FSS-2/EWT(1)/FS(v)-3/FCC/EWA(d)/EWA(h)    F      081-67    F    F      08	
TITLI eart SOUR	E: Irregular flows of high energy electrons dice th's radiation belts RCE: <u>Vsesoyuznaya konferentsiya po fizike kosmicheskogo prostranstva. Mosc</u> 5. Issledovaniya kosmicheskogo prostranstva (Space research); Trudy konfere	ow, intsii.
TOPI	IC TAGS: geomagnetic field, satellite data analysis, <u>see and "Elektron-1" and "Elektron-</u> STRACT: The authors analyze data obtained from " <u>Elektron-1" and "Elektron-</u> stract: The authors analyze data obtained from " <u>Elektron-1" and "Elektron-</u> stract: The authors analyze data obtained from " <u>Elektron-1" and "Elektron-</u> stract: The authors analyze data obtained from " <u>Elektron-1" and "Elektron-</u> stract: The authors analyze data obtained from " <u>Elektron-1" and "Elektron-</u> stract: The authors analyze data obtained from " <u>Elektron-1" and "Elektron-</u> stract: The authors analyze data obtained from " <u>Elektron-1" and "Elektron-</u> stract: The authors analyze data obtained from " <u>Elektron-1" and "Elektron-</u> stract: The authors analyze data obtained from " <u>Elektron-1" and "Elektron-</u> stract: <u>stract: stract</u> , <u>stract</u> , <u>strac</u>	
- lint	where first month of operation. The equation middle states the generation of the indicates the scribed. Analysis of data pertaining to the midnight meridian indicates the scribed. Analysis of data pertaining to the outer belt decreases by two tensity of the electrons at the boundary of the outer belt decreases. It is even orders of magnitude within a narrow range of radial distances. It is even orders of magnitude within a narrow range of the earth terminates on or shed that the radiation belt on the night side of the belt extends on the state of the belt extends of the belt extends of the belt extends of the belt extends on the state of the belt extends of the belt exten	uller i a
Ca	ard 1/2	

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ACCESSION NR: AT5023614		·
dipole approximation and e flows of electrons outside an increase in perturbation earth and at distances of	The L is the nominal McIlwain para expressed in earth radii.) It is the boundary of the earth's rad on of the geomagnetic field both $\sqrt{30,000}$ km from the earth. A th The experimental data support in the earth's magnetic field up a and 1 table.	liation belts appear with at the surface of the meoretical explanation is the hypothesis of a close
ASSOCIATION: none		
SUBMITTED: 02Sep65	ENCL: 00	SUB CODE: ES, SV
	OTHER: 010	ATD PRESS: 4105
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	L 3096-66 FSS-2/EWT(1)/FS(v)-3/FCC/EWA(d) TT/GS/GW ACCESSION NR: AT5023615 UR/0000/65/000/000/0433/0434 98	
1	AUTHORS: Varnov, S. N.; Chudakov, A. Ye,; Vakulov, P. V.; Gorchakov, Ye. V. B+1	1
1	AUTHORS: <u>Vernov, S. N.;</u> Chudakov, A. Ye,; <u>Vakulov, P. V.;</u> Gorchakov, Ye. V. B+/ Logachev, Yu. I.; Nikolayev, A. G.; Rubinshteyn, I. A.; Sosnovets, E. N.; <sup>49,55</sup>	1.
	Ternovskaya, M. V. 49.55 44.25 44.55	·
	TITLE: Pulsations of the earth's magnetic field, from the measurements taken by the Elektron-3 satellite	
.	SOURCE: Vsesoyuznava konferentsiya po fizike kosmicheskogo prostranstva, Moscow,	
Ī	<u>1965.</u> Issledovaniya kosmicheskogo prostranstva (Space research); trudy konferentsii. Moscow, Izd-vo Nauka, 1965, 433-434	
	TOPIC TAGS: satellite, satellite data analysis, pulse counter, pulse amplifier, pulse amplitude, <u>earth magnetic field</u>	
	ABSTRACT: The Elektron-3 satellite, launched on July 11, 1964, carried a coil with ferrite core. Signals from this coil were transmitted to two amplifying circuits,	8
	one for the band of 1-10 cps, the other for 30-300 cps. Both circuits recorded pulses with amplitudes exceeding $\sqrt{1}$ , $\sqrt{5}$ , $\sqrt{25}$ Y. The type and operation of	
	the memory bank are briefly described. From a small amount of data processed it can be seen that no pulses with the amplitudes $\gtrsim 25$ Y were recorded, that at	
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ACCESSION NE	R: AT5023615			ne van de lan ne verse en ser ar ger ook en op ook op op ook op	•••••••••••••••••••••••••••••••••••••••	
by the low-f noted that t generally gr	the intermediate requency circuit the number of mag ceater in the fre	l γ) the count e e sensitivity (> t and about 1 by gnetic field puls equency region of	$5\gamma$ about 2- the high-frequences with the amplitude the second	3 pulses were ency circuit. plitude よう n in the regi	e recorded It is Y is	
ASSOCIATION:	none	recorded by the 1	Crease in some .cw-frequency c:	lrcuit but no	t by the [04]	
ASSOCIATION:	none character	recorded by the 1	ow-frequency c:	lrcuit but no	t by the [04]	
ASSOCIATION:	none 02Sep65	recorded by the 1 Confluence Light W.F. L. J. doc.	ow-frequency c:	SUB CODE:	t by the [04] anstva <b>ES, SV</b>	
Normally, th high-frequen ASSOCIATION:	none 02Sep65	recorded by the 1 conference of the 1 conferen	ow-frequency c:	ircuit but no	t by the [04] anstva <b>ES, SV</b>	
Normally, th high-frequen ASSOCIATION:	none 02Sep65	recorded by the 1 conference of the 1 conferen	ow-frequency c:	SUB CODE:	t by the [04] anstva <b>ES, SV</b>	

ACCESSION NR: AT5023628	UR/0000/65/000/000/0502/0506
UTHOR: <u>Vernov, S. L</u> ; <u>Vakulov, P. V.;</u> Zat <u>kholopkov, V. P.;</u> <u>Chudakov, A. Ye.</u> 44,55 TTLE: Primary cosmic radiation investigat	
OURCE: <u>Vsesoyuznaya konferentsiya po fizi</u> 965. Issledovaniya kosmicheskogo prostran oscow, Izd-vo Nauka, 1965, 502-506	ke kosmi chaskopp another with 55
DPIC TAGS: cosmic ray, cosmic radiation, adiation, <u>Elektron 2</u> , <u>Elektron 4</u> 12,55	12,44,23,
BSTRACT: Experimental data obtained by Ela Lon are presented and interpreted. The data November 1964, were obtained primarily by a average frequency of 20 pulses/sec. The seping them outside the earth's radiation hourt frequency as the thickness of the scree onclude that the purchase radiation to be a screen by the screen state the screen state the screen scients.	ta, covering the period 30 January to means of gas-discharge counters with apogee of the satellites was 68,000 km, belts most of the time. The higher
onclude that the primary radiation did not 10 Mev range. Two types of radiation inter d 1/3	contain newbialas withhis she to se